# Testing theories of gravity with X-ray reflection spectroscopy

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• Why test general relativity?

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- Alternatives to general relativity.

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- Alternatives to general relativity.
- Black holes as the best tests.
- Alternative black holes.
- What about gravitational waves?

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- Within GR, described by the Kerr metric.
- Parameters in the Kerr metric: M, a, |a| < M.
- Parametrically deformed in alternative theories.
- Top-down and bottom-up approaches.

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# Accreting black holes



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## Radiation profile



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#### Broadening of a line



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#### • Kerr metric:

$$ds^{2} = -\left(1 - \frac{2Mr}{\Sigma}\right)dt^{2} + \frac{\Sigma}{\Delta}dr^{2} - \frac{2Mar\sin^{2}\theta}{\Sigma}dt\,d\phi$$
$$+ \Sigma d\theta^{2} + \left(r^{2} + a^{2} + \frac{2Ma^{2}r\sin^{2}\theta}{\Sigma}\sin^{2}\theta\right)d\phi^{2} \quad (1)$$
$$\Delta \equiv r^{2} - 2Mr + a^{2}, \quad \Sigma \equiv r^{2} + a^{2}\cos^{2}\theta \quad (2)$$

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#### Iron lines within GR



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$$\mathcal{L} = \frac{1}{2}R - \frac{1}{4}\partial_{\mu}\partial^{\mu}\phi + \frac{\alpha'}{8g^2}e^{\phi}\left(R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} - 4R_{\mu\nu}R^{\mu\nu} + R^2\right)$$

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#### Black holes beyond GR: bottom-up

• Johannsen metric<sup>1</sup>:

$$ds^{2} = -\frac{\tilde{\Sigma}[\Delta - a^{2}A_{2}^{2}\sin^{2}\theta]}{\Xi}dt^{2} + \frac{\tilde{\Sigma}}{\Delta A_{5}}dr^{2}$$
$$-\frac{a[(r^{2} + a^{2})A_{1}A_{2} - \Delta]}{\Xi}dt\,d\phi + \tilde{\Sigma}d\theta^{2}$$
$$+\frac{\tilde{\Sigma}\sin^{2}\theta[(r^{2} + a^{2})^{2}A_{1}^{2} - a^{2}\Delta\sin^{2}\theta]}{\Xi}d\phi^{2} \qquad (3)$$

$$\Xi = ((r^2 + a^2)A_1 - a^2A_2\sin^2\theta)^2,$$
  

$$A_1 = 1 + \alpha_{13}(M/r)^3 + \dots, \quad A_2 = 1 + \alpha_{22}(M/r)^2 + \dots$$
  

$$A_5 = 1 + \alpha_{52}(M/r)^2 + \dots, \quad \tilde{\Sigma} = \Sigma + \epsilon_3(M^3/r) + \dots$$

1 Johannsen, PRD, 88, 044002 (2013)

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#### Iron lines beyond GR



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| Model | relxi | 11_nk<1> Sou | urce No.: 1 | Active | /0n      |         |
|-------|-------|--------------|-------------|--------|----------|---------|
| Model | Model | Component    | Parameter   | Unit   | Value    |         |
| par   | comp  |              |             |        |          |         |
| 1     | 1     | relxill_nk   | Index1      |        | 3.00000  | frozen  |
| 2     | 1     | relxill_nk   | Index2      |        | 3.00000  | frozen  |
| 3     | 1     | relxill_nk   | Rbr         |        | 15.0000  | frozen  |
| 4     | 1     | relxill_nk   | а           |        | 0.998000 | +/- 0.0 |
| 5     | 1     | relxill_nk   | Incl        | deg    | 30.0000  | +/- 0.0 |
| 6     | 1     | relxill_nk   | Rin         |        | -1.00000 | frozen  |
| 7     | 1     | relxill_nk   | Rout        |        | 400.000  | frozen  |
| 8     | 1     | relxill_nk   | z           |        | 0.0      | frozen  |
| 9     | 1     | relxill_nk   | gamma       |        | 2.00000  | +/- 0.0 |
| 10    | 1     | relxill_nk   | logxi       |        | 3.10000  | +/- 0.0 |
| 11    | 1     | relxill_nk   | Afe         |        | 5.00000  | +/- 0.0 |
| 12    | 1     | relxill_nk   | Ecut        | keV    | 300.000  | frozen  |
| 13    | 1     | relxill_nk   | refl_frac   |        | 3.00000  | +/- 0.0 |
| 14    | 1     | relxill_nk   | defpar_type | e      | 1.00000  | frozen  |
| 15    | 1     | relxill_nk   | defpar_valu | ue     | 0.0      | +/- 0.0 |
| 16    | 1     | relxill_nk   | norm        |        | 1.00000  | +/- 0.0 |
|       |       |              |             |        |          |         |

#### Constraints with current instruments



arXiv:1804.10380, under review

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#### Scope with future instruments



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#### A glimpse into the future



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#### A glimpse into the future



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#### The curious case of Einstein dilaton Gauss Bonnet



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arXiv:1803.10819, to appear in PLB

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#### Thank you!

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## **ISCO** contours



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