

# *An Excess of Dusty Starbursts at $z=2.2$*

*Helmut Dannerbauer (Vienna)*



**Kurk, De Breuck, Altieri, Coia, Emonts, Galametz, Hatch, Kodama, Koyama, Miley, Röttgering, Seymour, Tanaka, Sanchez-Portal, Santos, Valtchanov, Venemans, Ziegler**

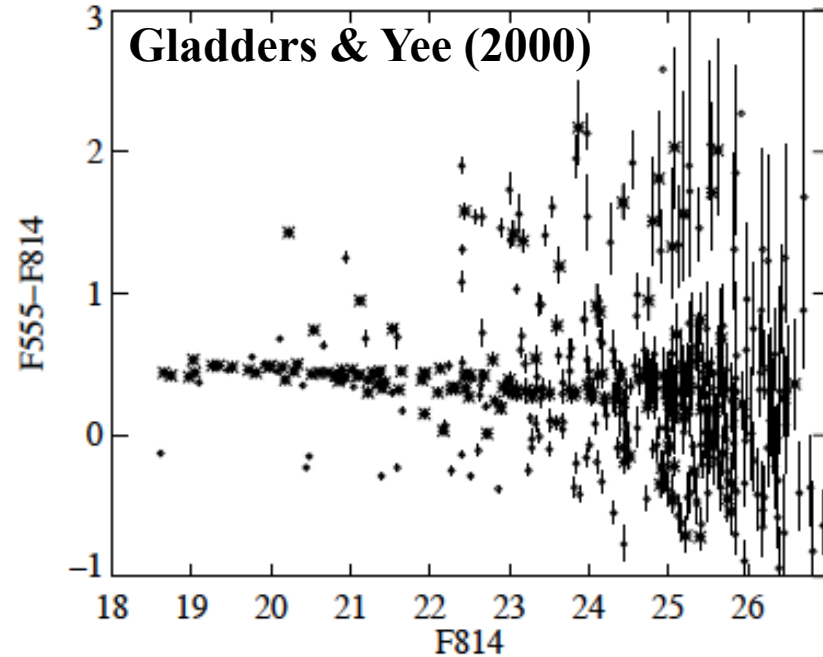
**→ *Dannerbauer, Kurk, De Breuck et al., 2014, A&A, 2014, 570, 55***

**→ *Dannerbauer, Emonts et al., 2015, to be submitted***

# Motivation

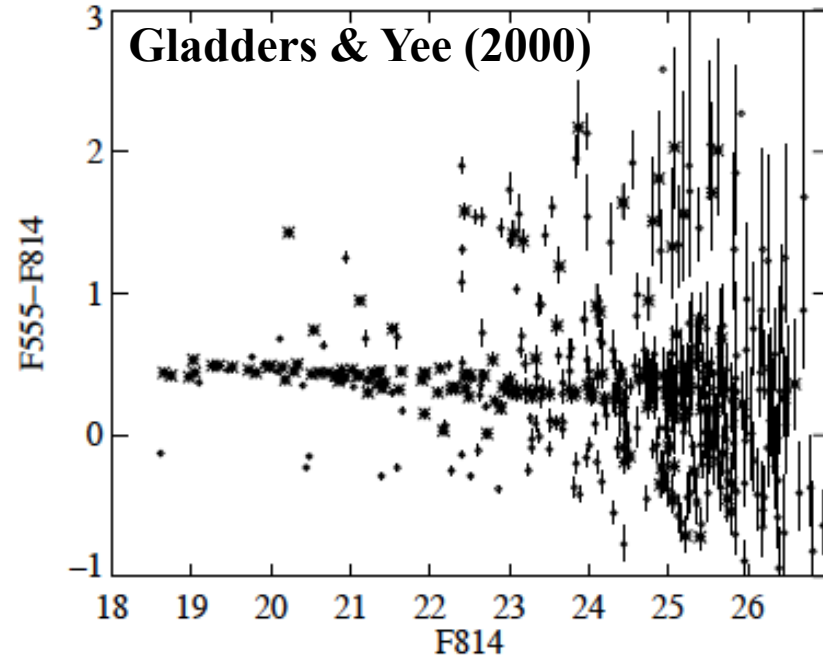
formation red sequence

complete census (pop., SFR) of overdensities

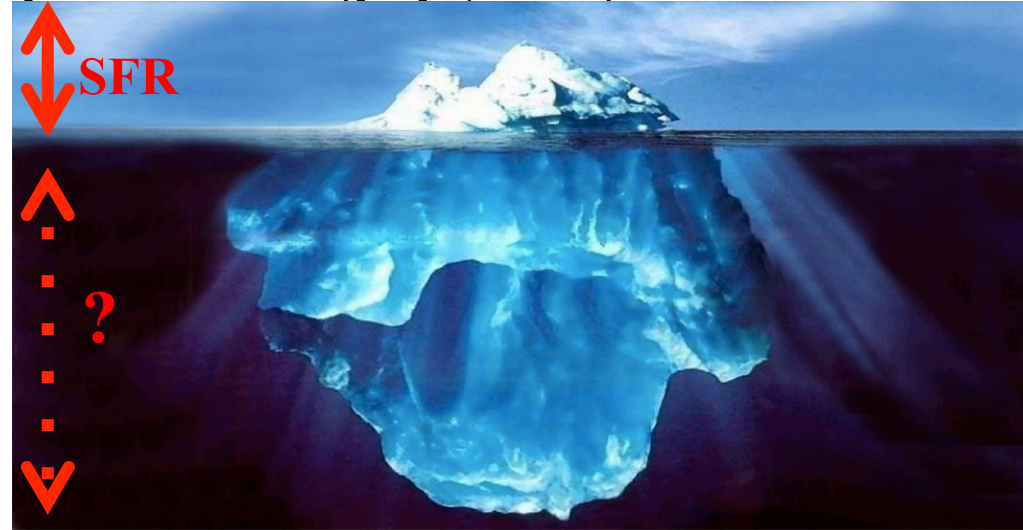


# Motivation

formation red sequence



complete census (pop., SFR) of overdensities



**→ search for massive galaxies in formation**



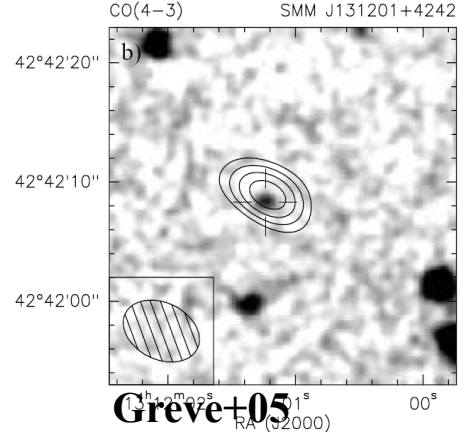
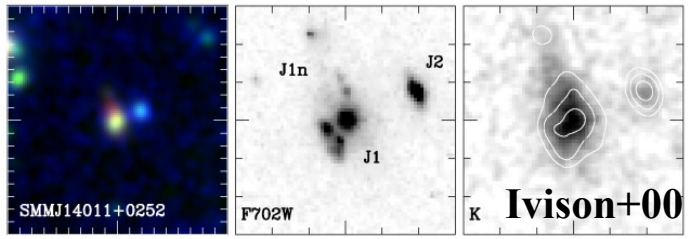
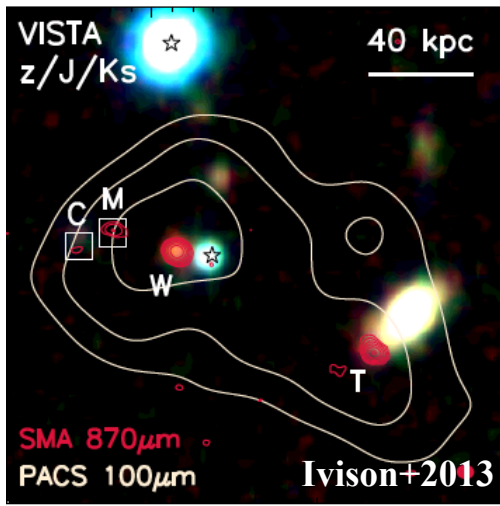
# Detecting Large Scale Structure in the (Sub)mm

## Submm-selected galaxies (SMGs):

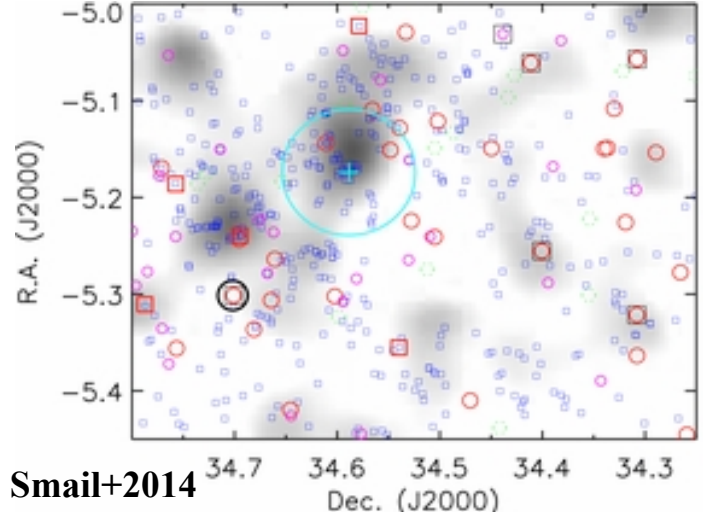
- very massive up to  $10^{11} M_{\odot}$
- gas-rich
- high SFR: several  $100 M_{\odot}/\text{yr}$
- merger-like morphology
- ellipticals in formation

→ *excellent tracers of mass-density peaks*

## SMGs within overdensities



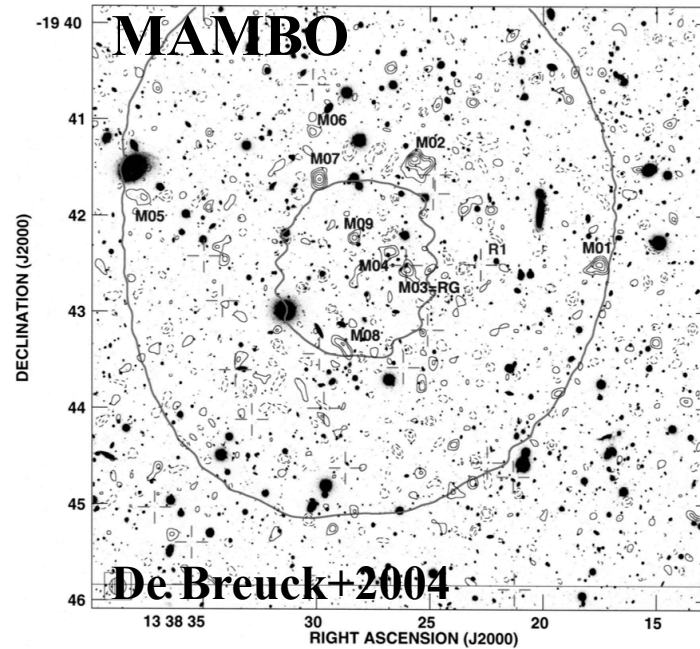
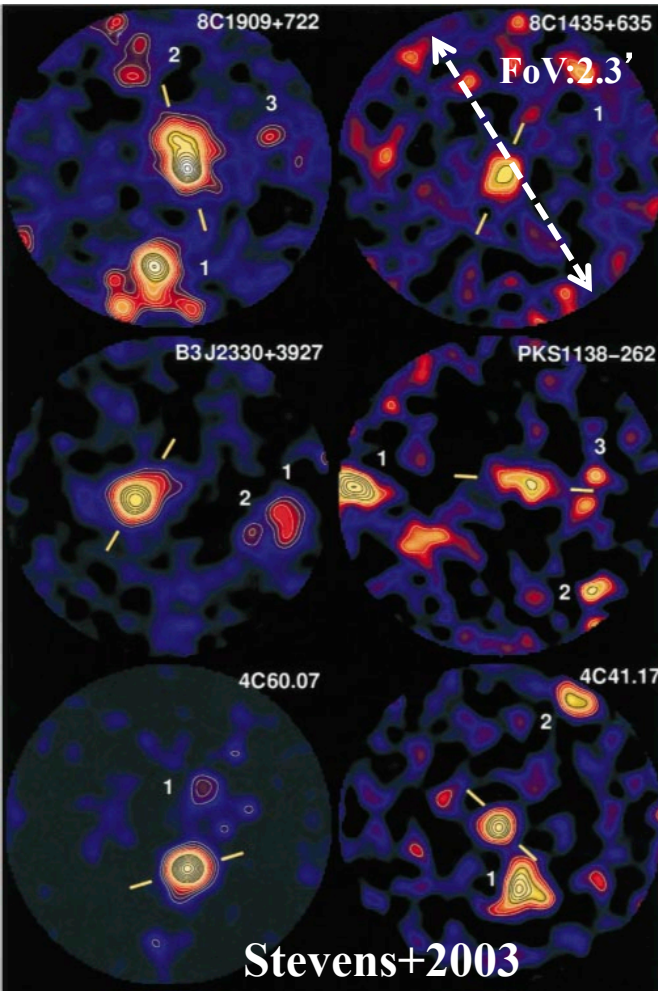
## search for overdensities in the (sub)mm



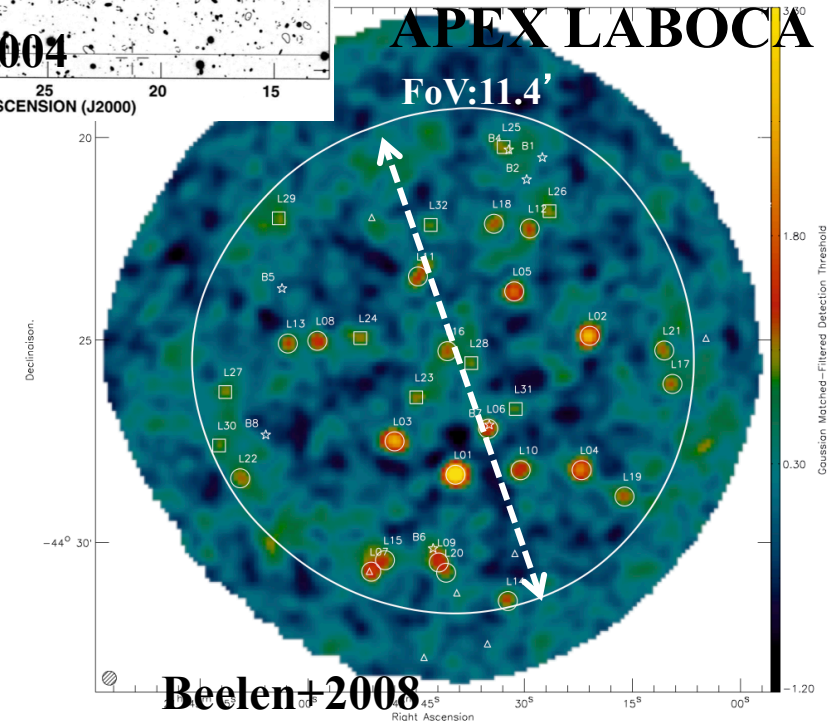


# Previous Work

## SCUBA

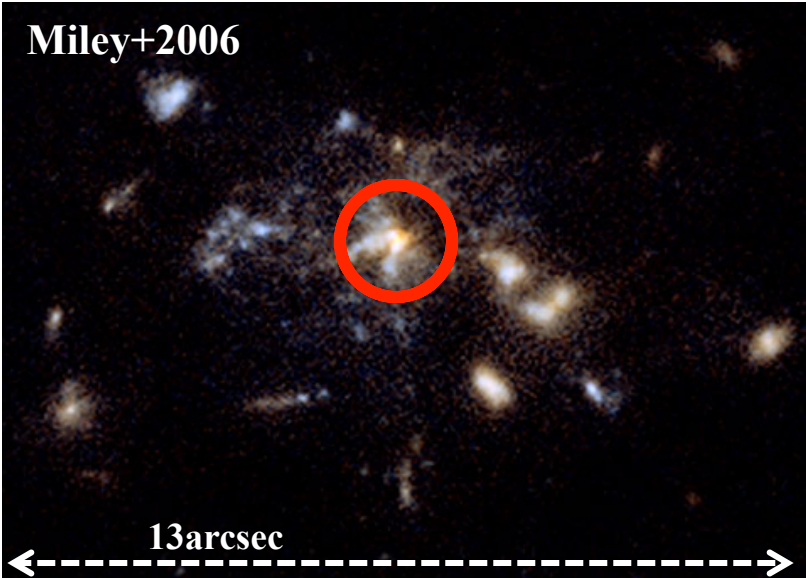


## APEX LABOCA



# Protocluster MRC1138 @ $z=2.16$

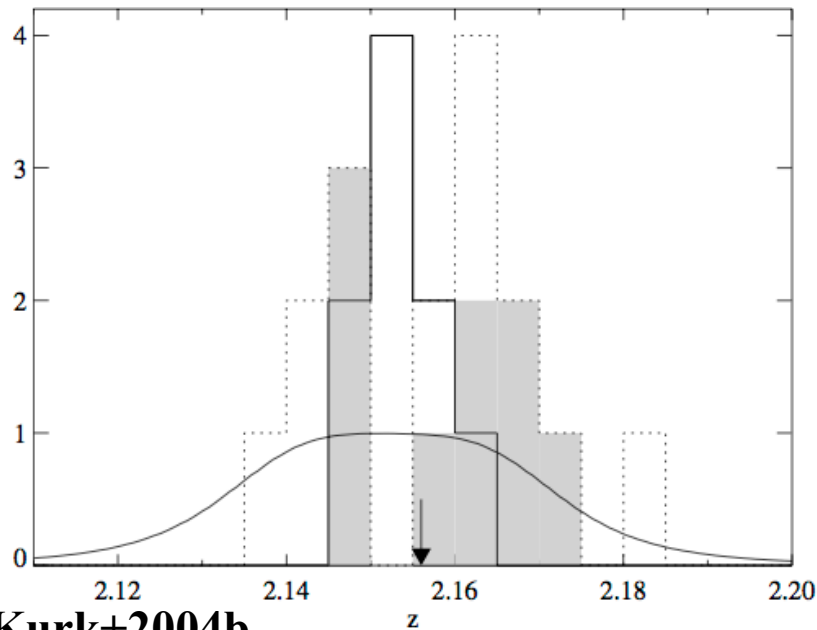
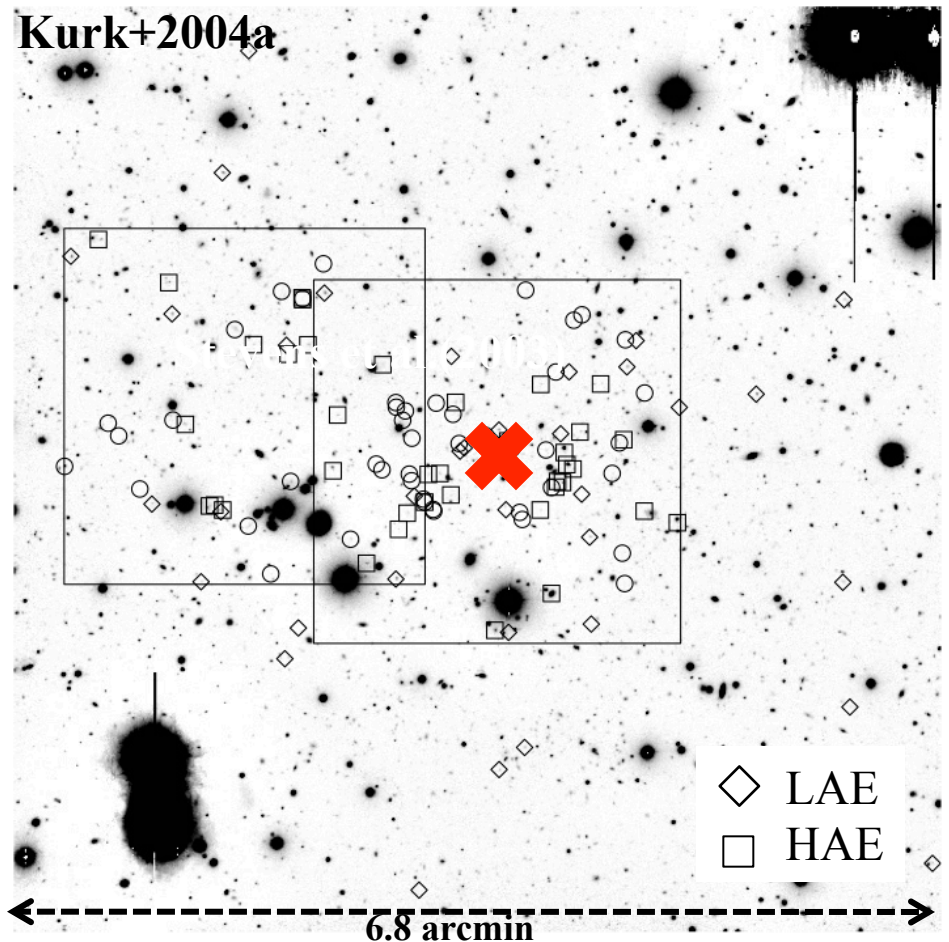
Miley+2006



Protocluster:

structure/collection of galaxies - not virialized -, which will evolve into a galaxy cluster @  $z=0$

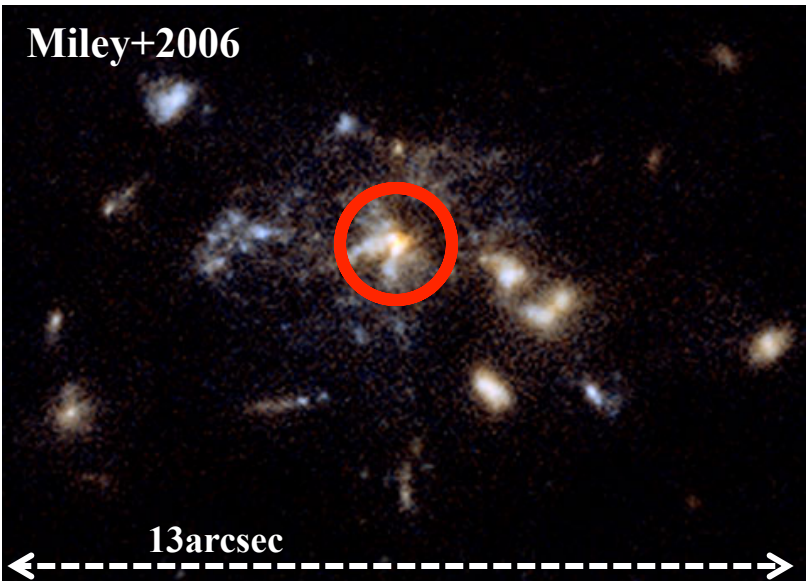
Kurk+2004a



Kurk+2004b

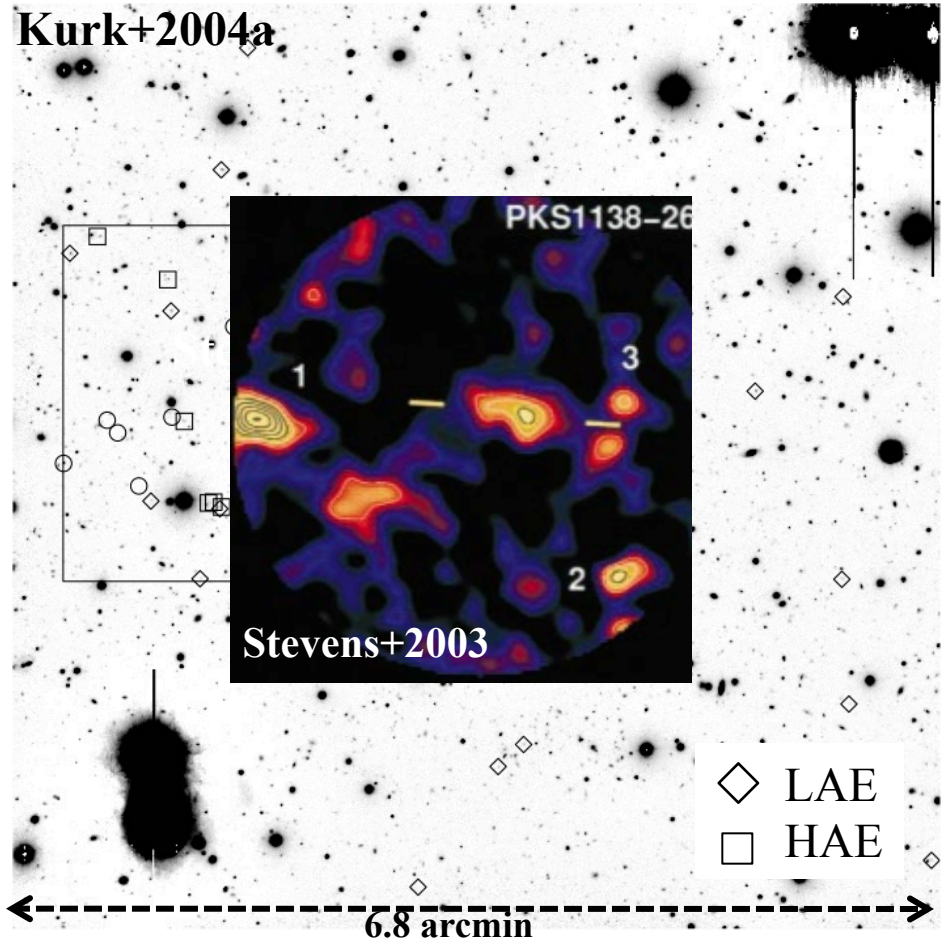
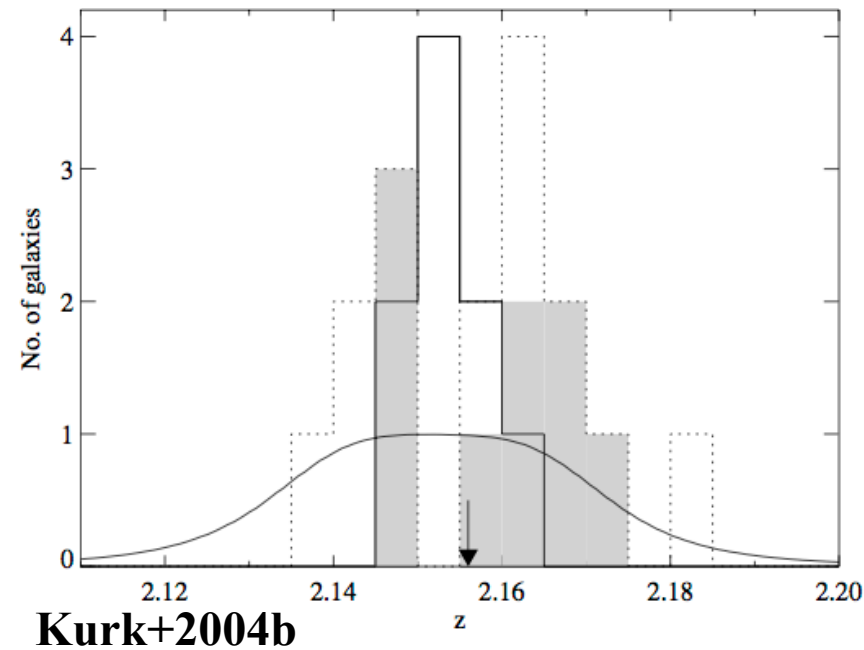


# Protocluster MRC1138 @ $z=2.16$



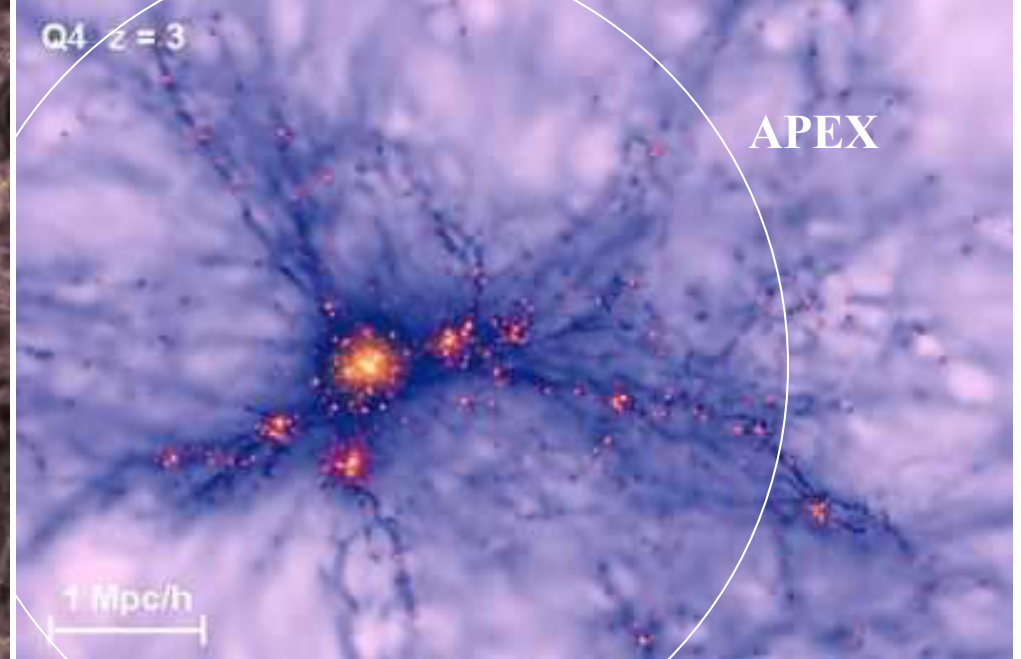
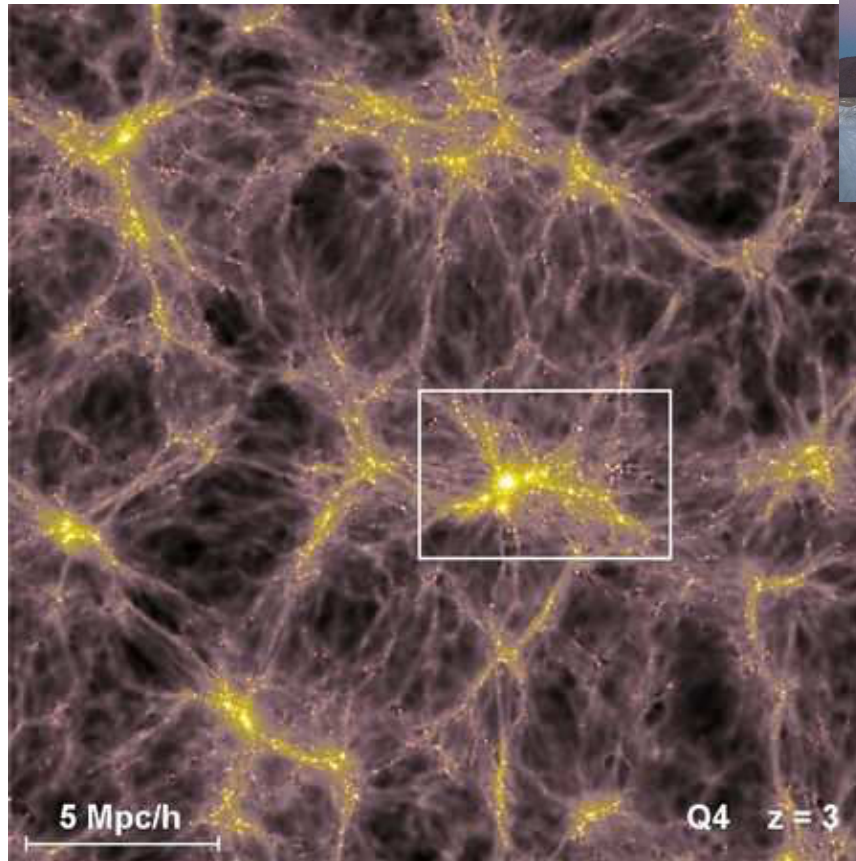
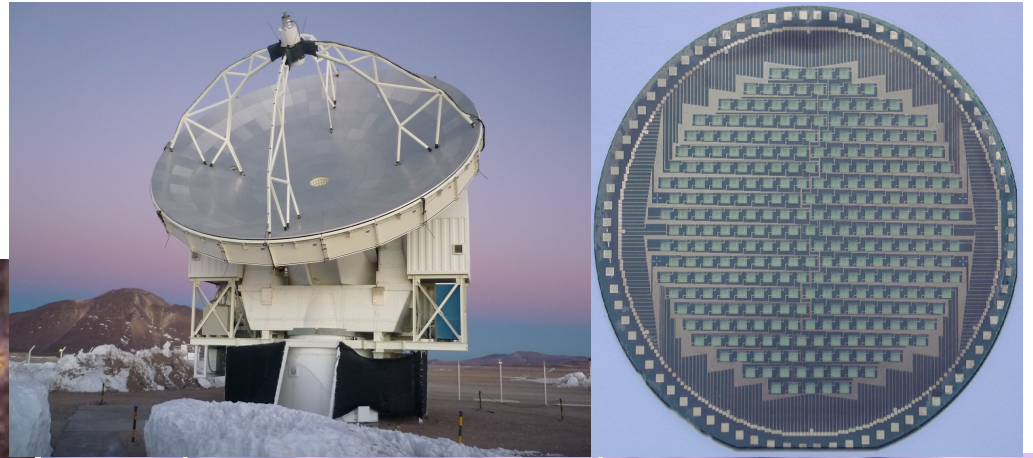
## Protocluster:

structure/collection of galaxies - not virialized -, which will evolve into a galaxy cluster @  $z=0$





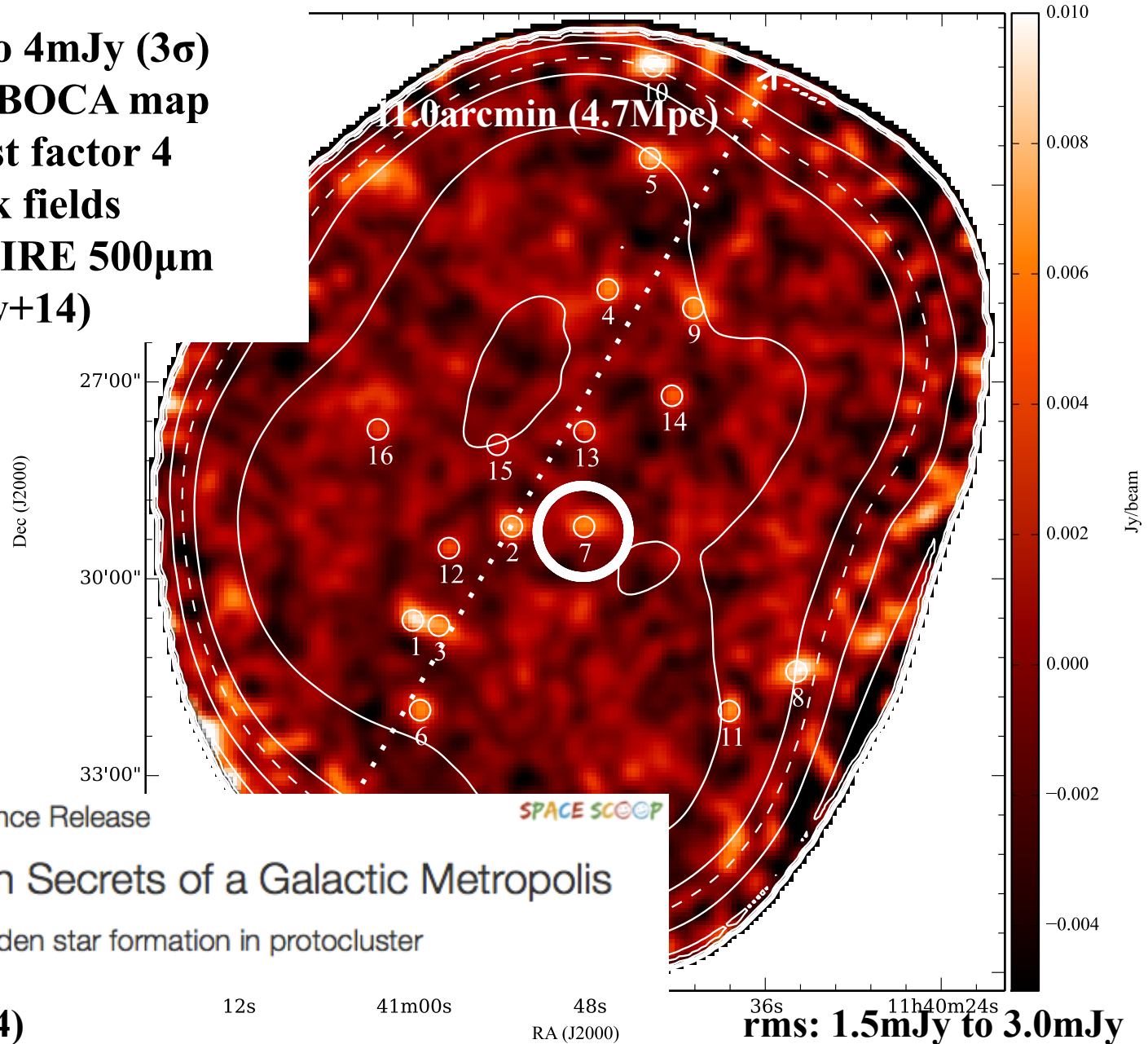
# APEX LABOCA Observations



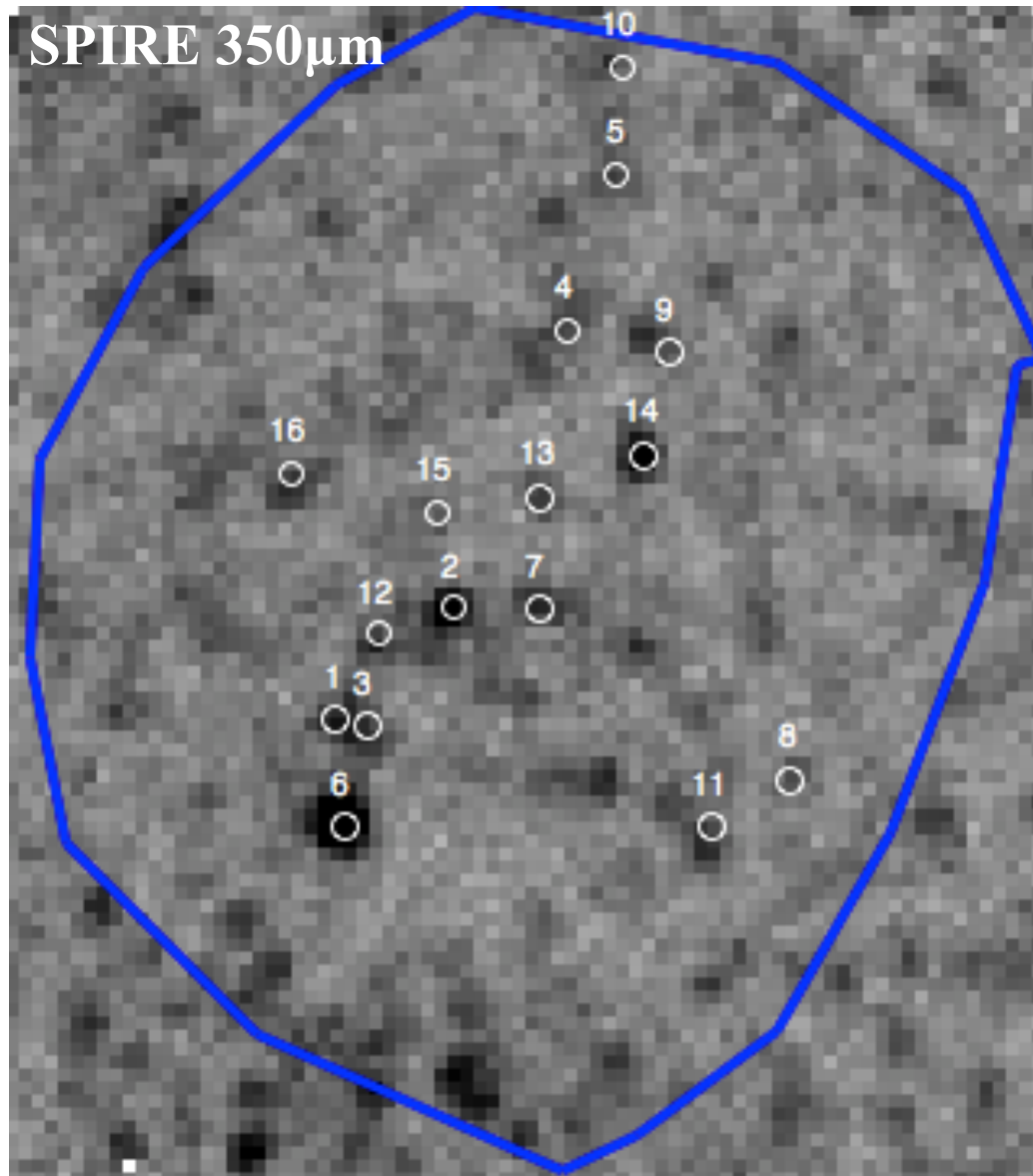
Springel & Hernquist (2003)

# APEX LABOCA Observations

- 16 sources down to 4mJy ( $3\sigma$ )
- one of deepest LABOCA map
- overdensity at least factor 4 compared to blank fields
- consistent with SPIRE 500 $\mu$ m overdensity (Rigby+14)



# Source Reliability - Herschel Counterparts



Vatlchanov+2013; Rigby+2014

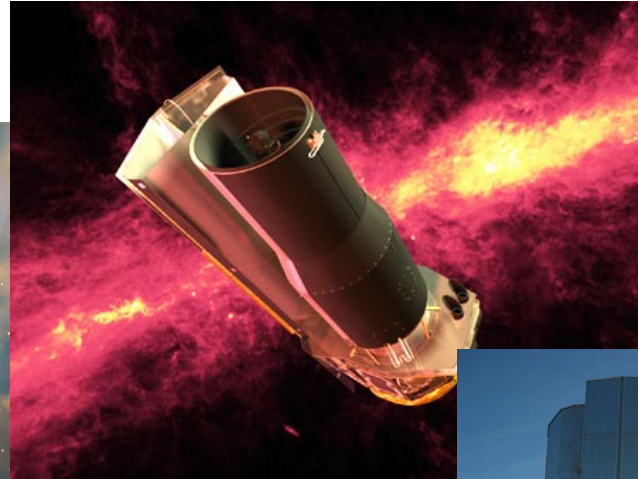
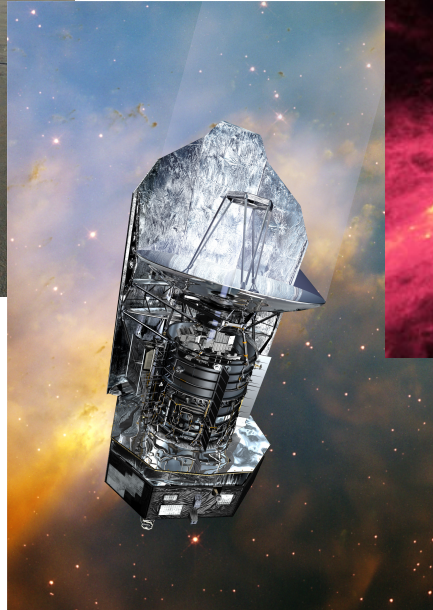


# Identification Work

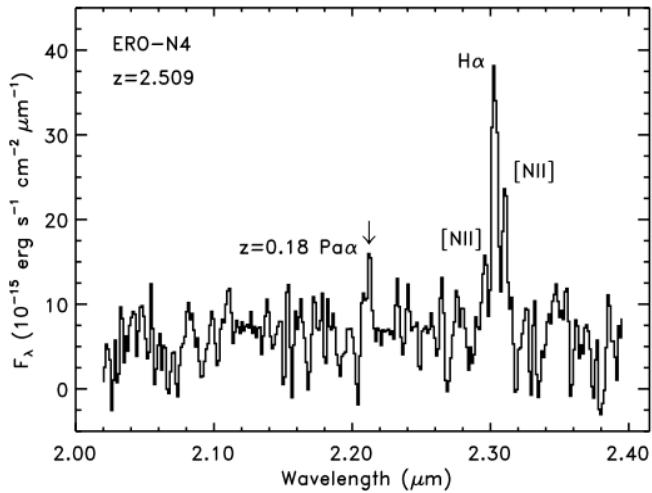
- best approach would be mm-interferometry



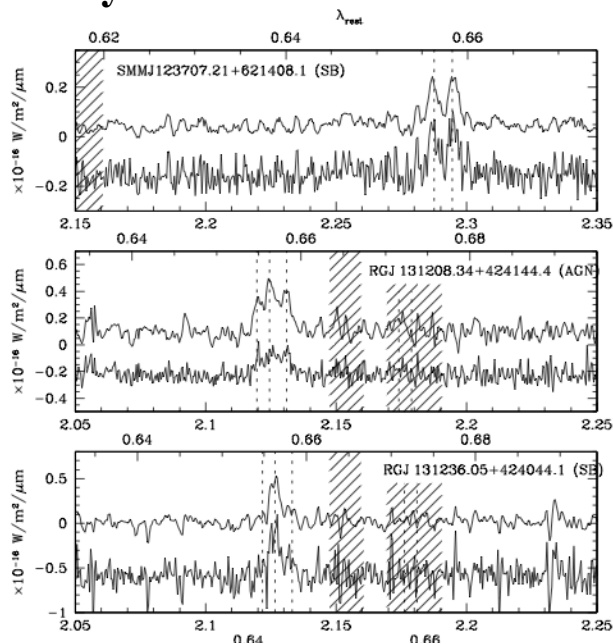
- counterpart identification could be done based on probabilistic statistic  
→corrected Poissonian Probability (p-statistics; Downes+86)



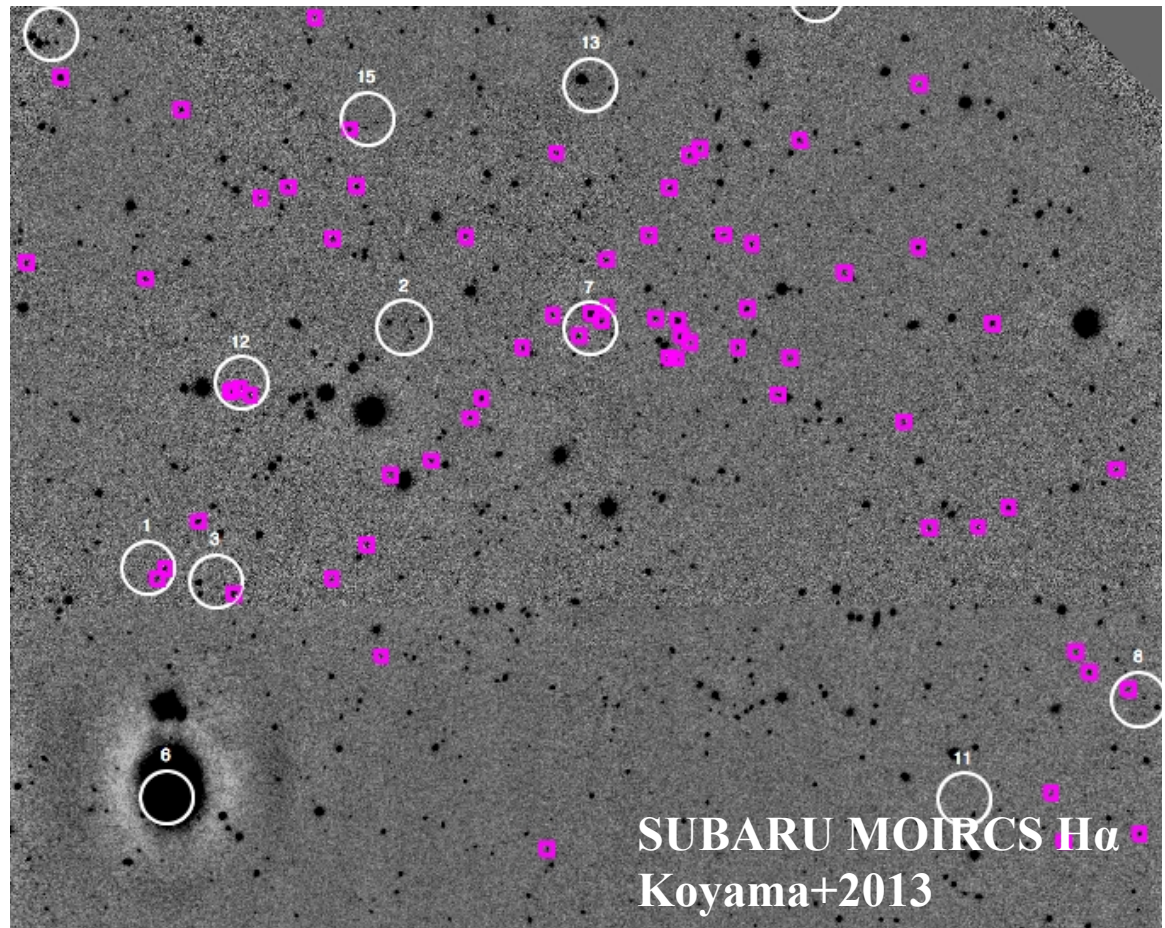
# HAEs associated with SMGs



Frayer+2004

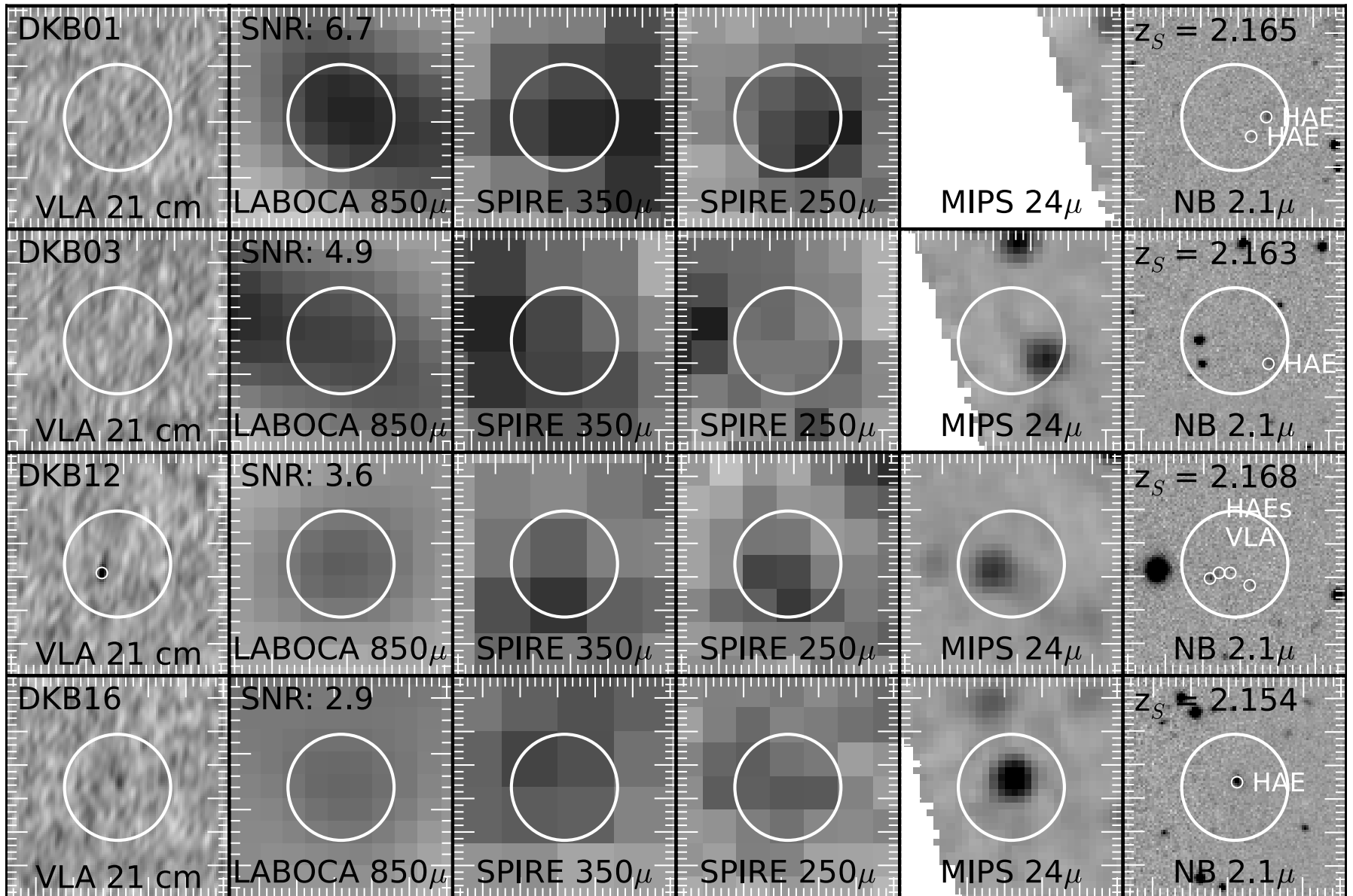


Swinbank+2003



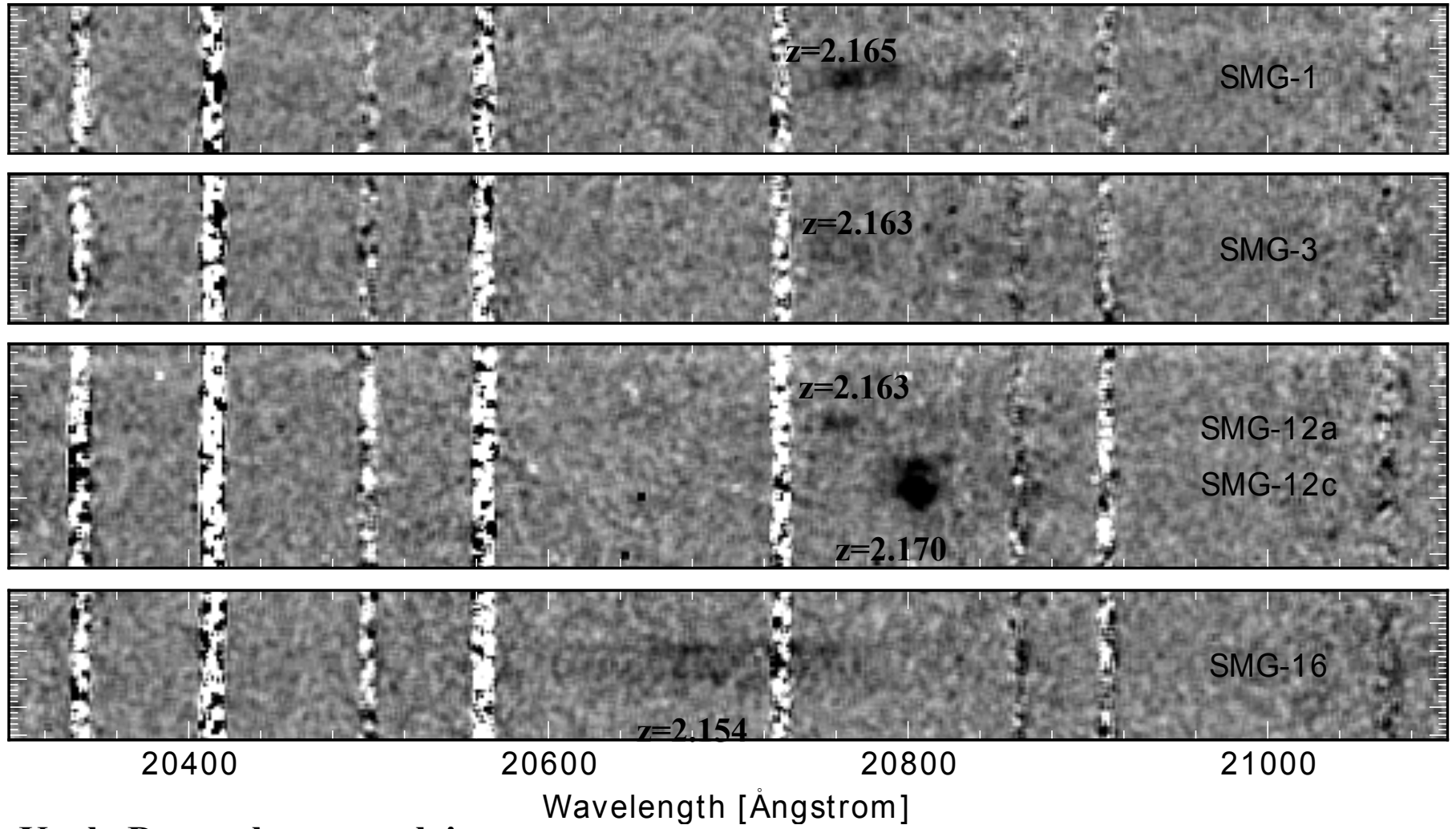
applying p-statistics on HAEs

# Counterparts



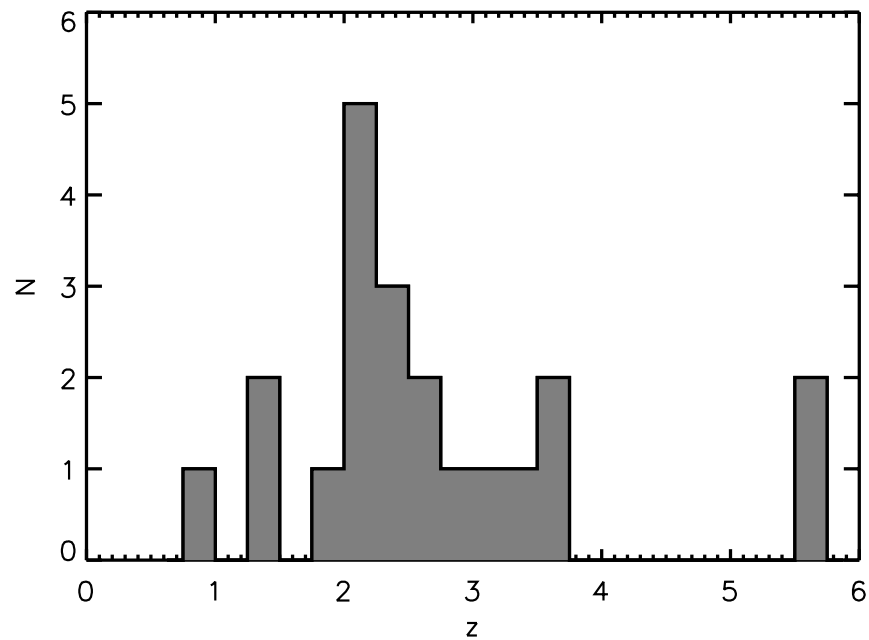
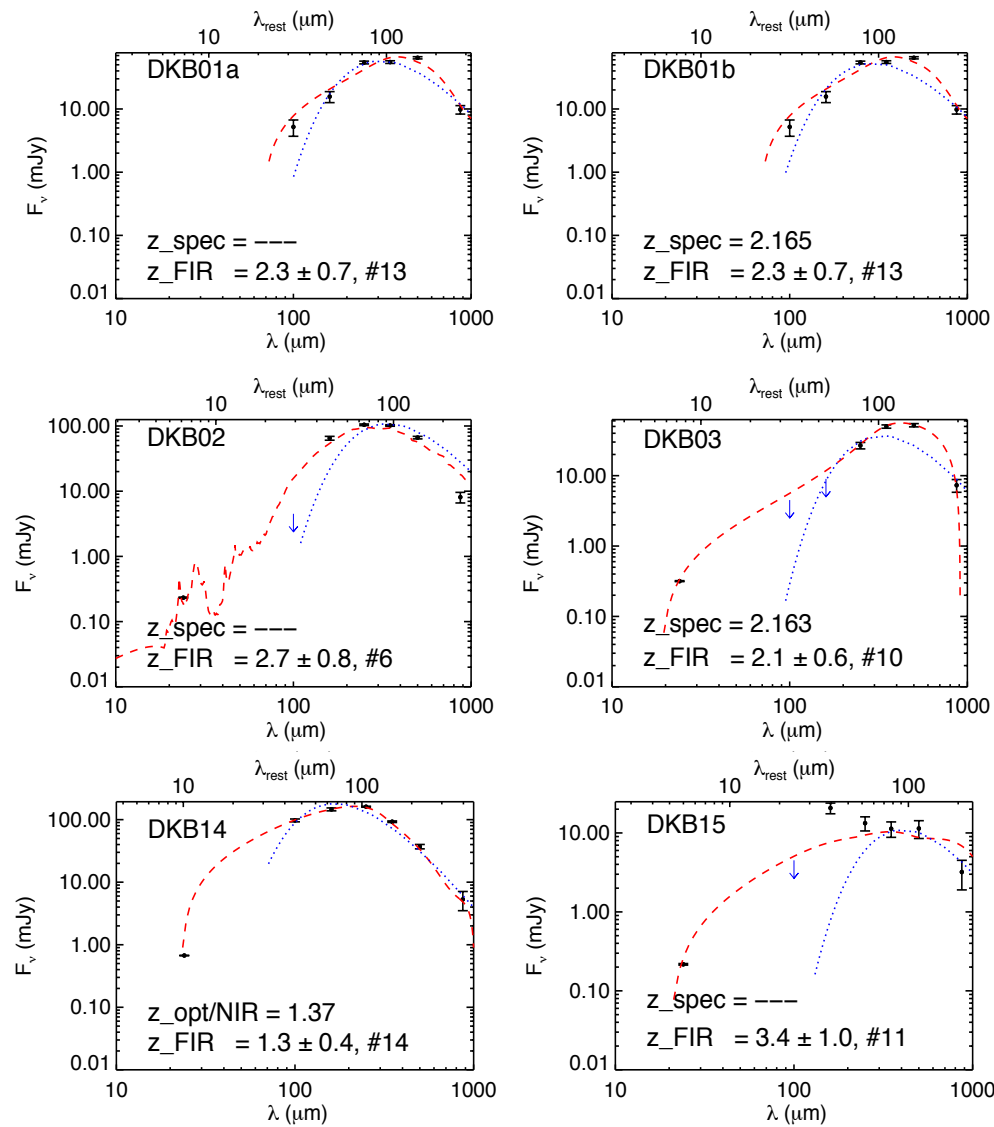


# ISAAC Spectroscopy

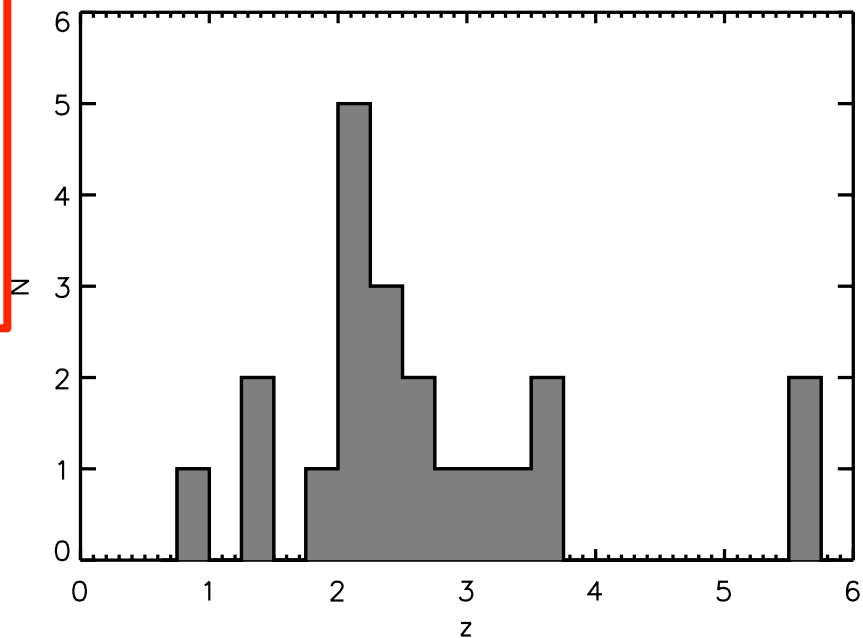
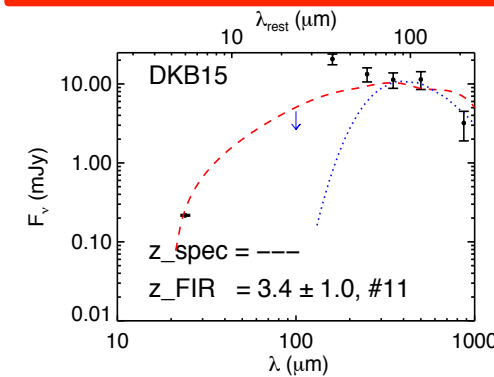
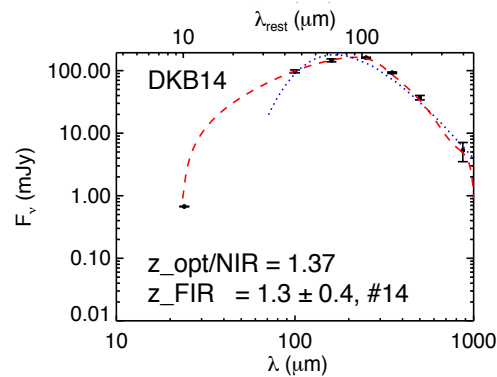
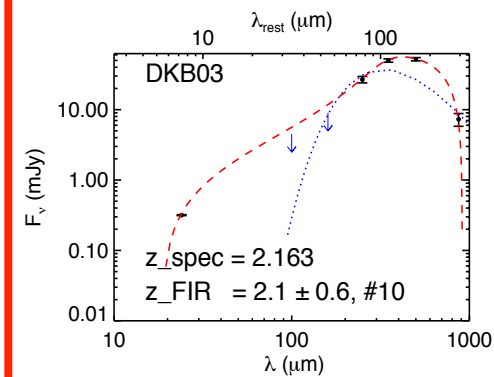
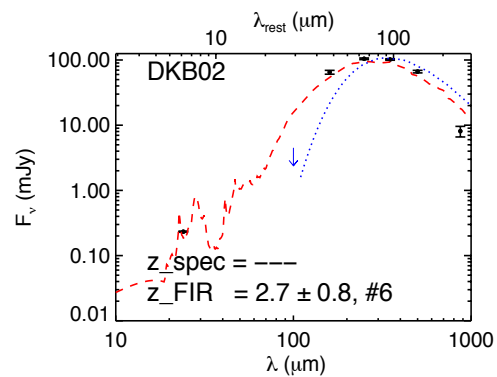
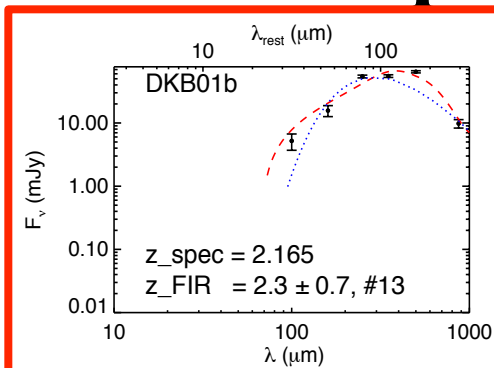
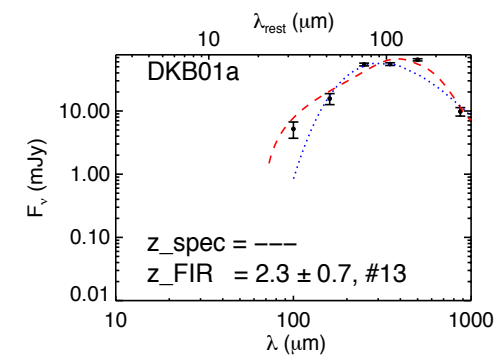


Kurk, Dannerbauer et al. in prep

# FIR-photo-z



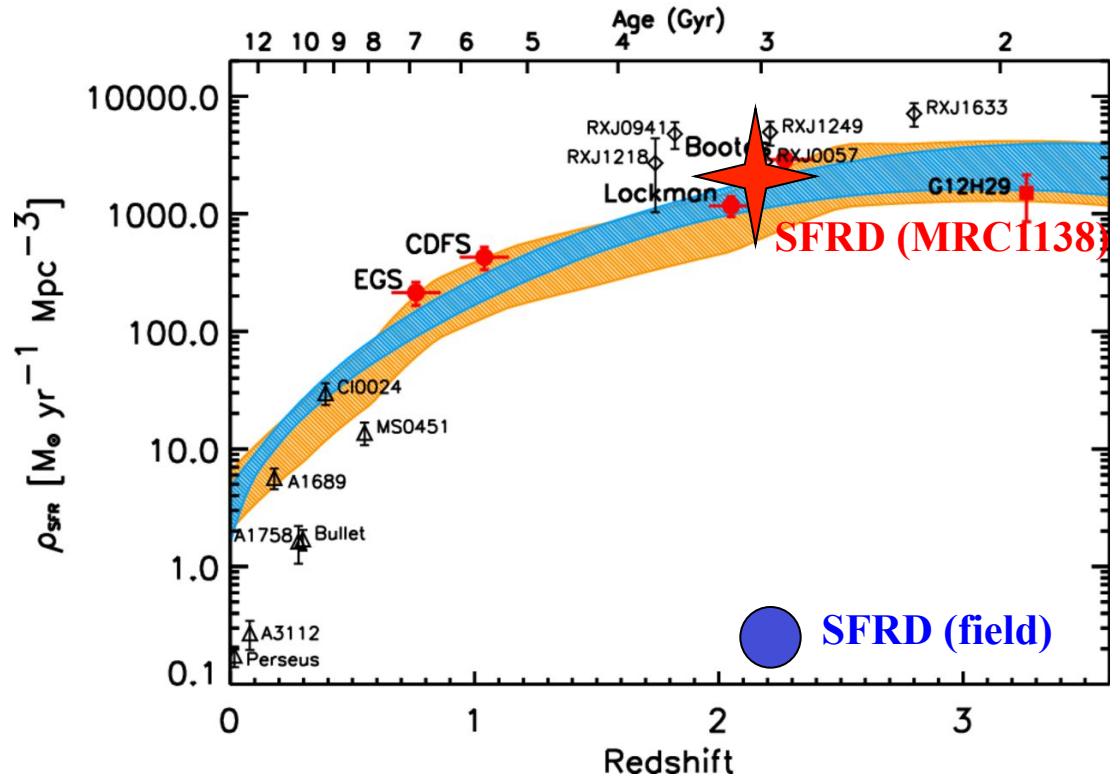
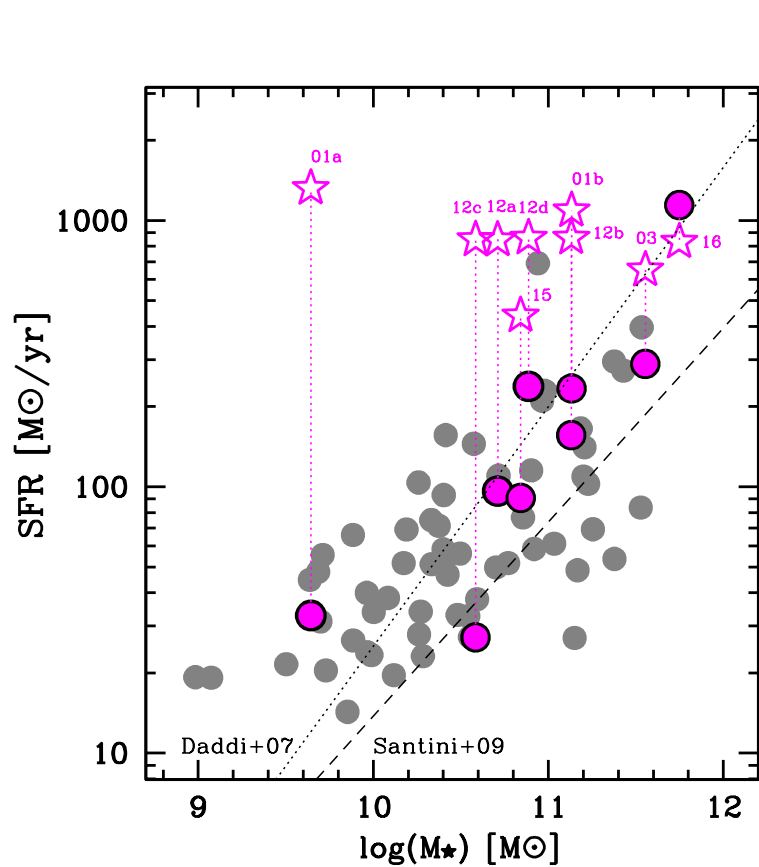
# FIR-photo-z



**ISAAC sources @ $z=2.16$ :**  
 **$z_{\text{phot\_IR}}$  consistent with  $z_{\text{spec\_H}\alpha}$**



# Dust Obscuration and Star Formation

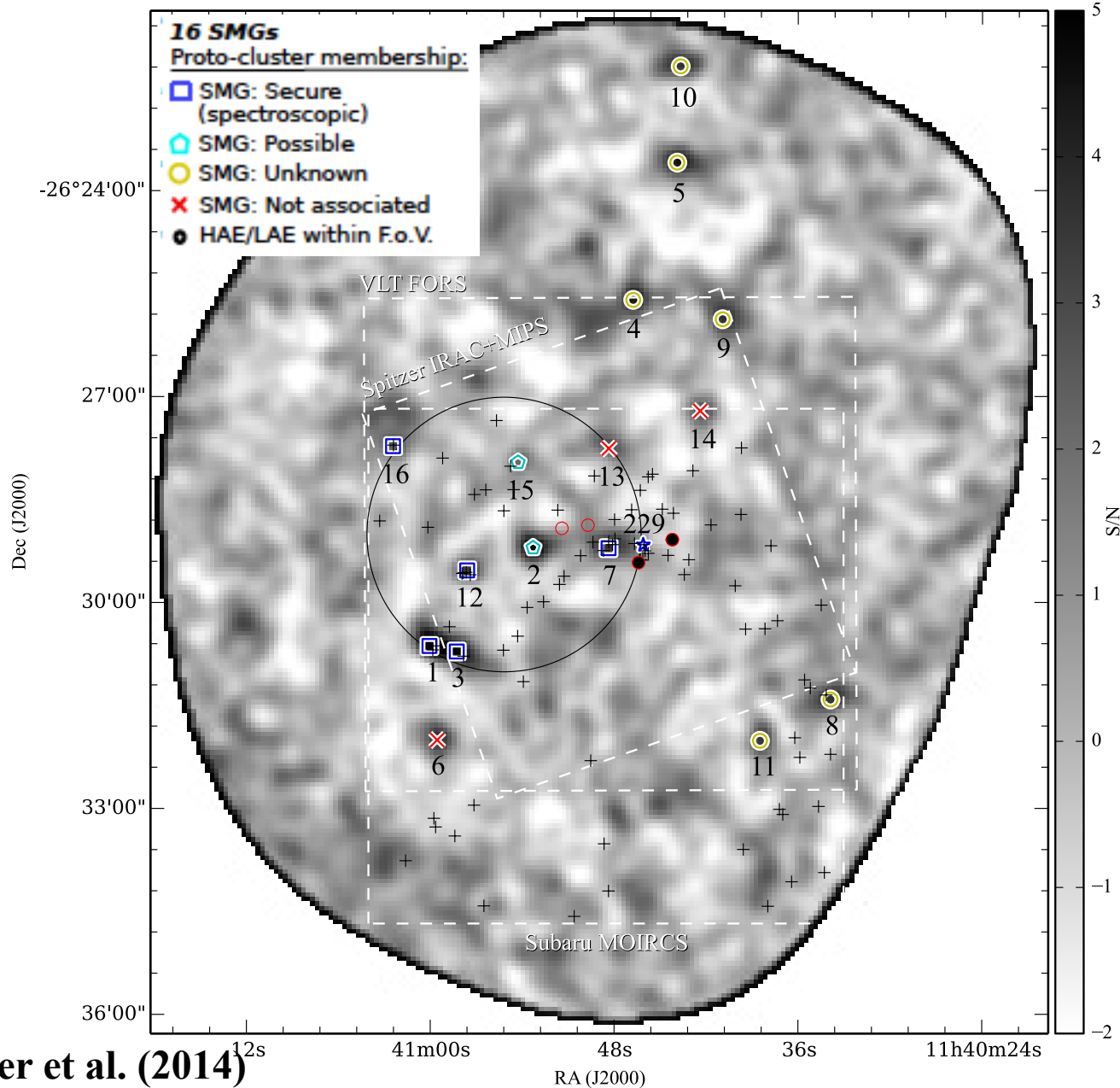


Clements+2014

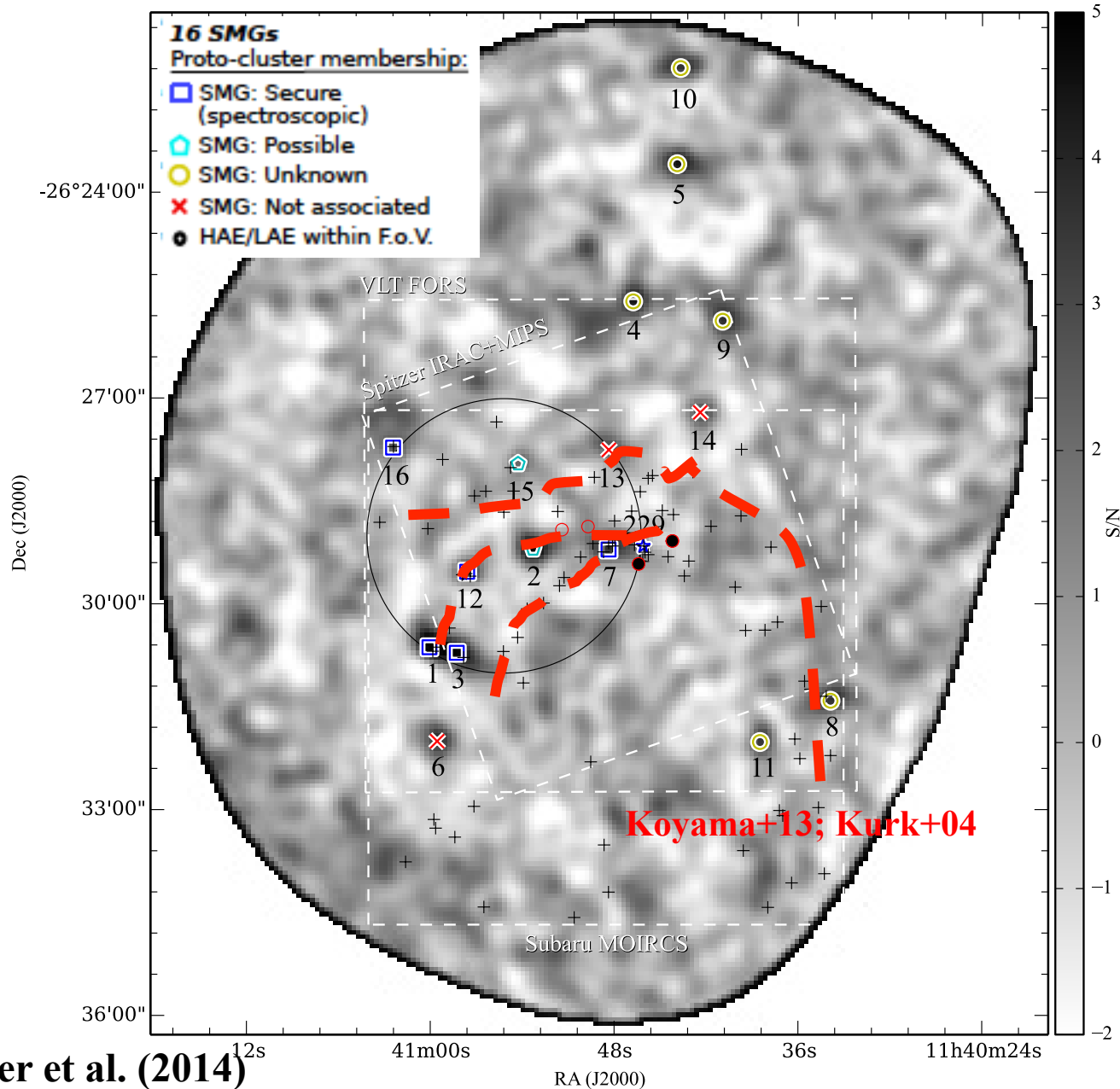
$\text{SFR}(\text{IR}) = 5\text{-}10 \times \text{SFR}(\text{H}\alpha)$

≠ simulations by Granato+15  
 underpredict by a factor 5 to 10 the  
 SFRD → revise feedback/SF models

# Proto-Cluster Membership

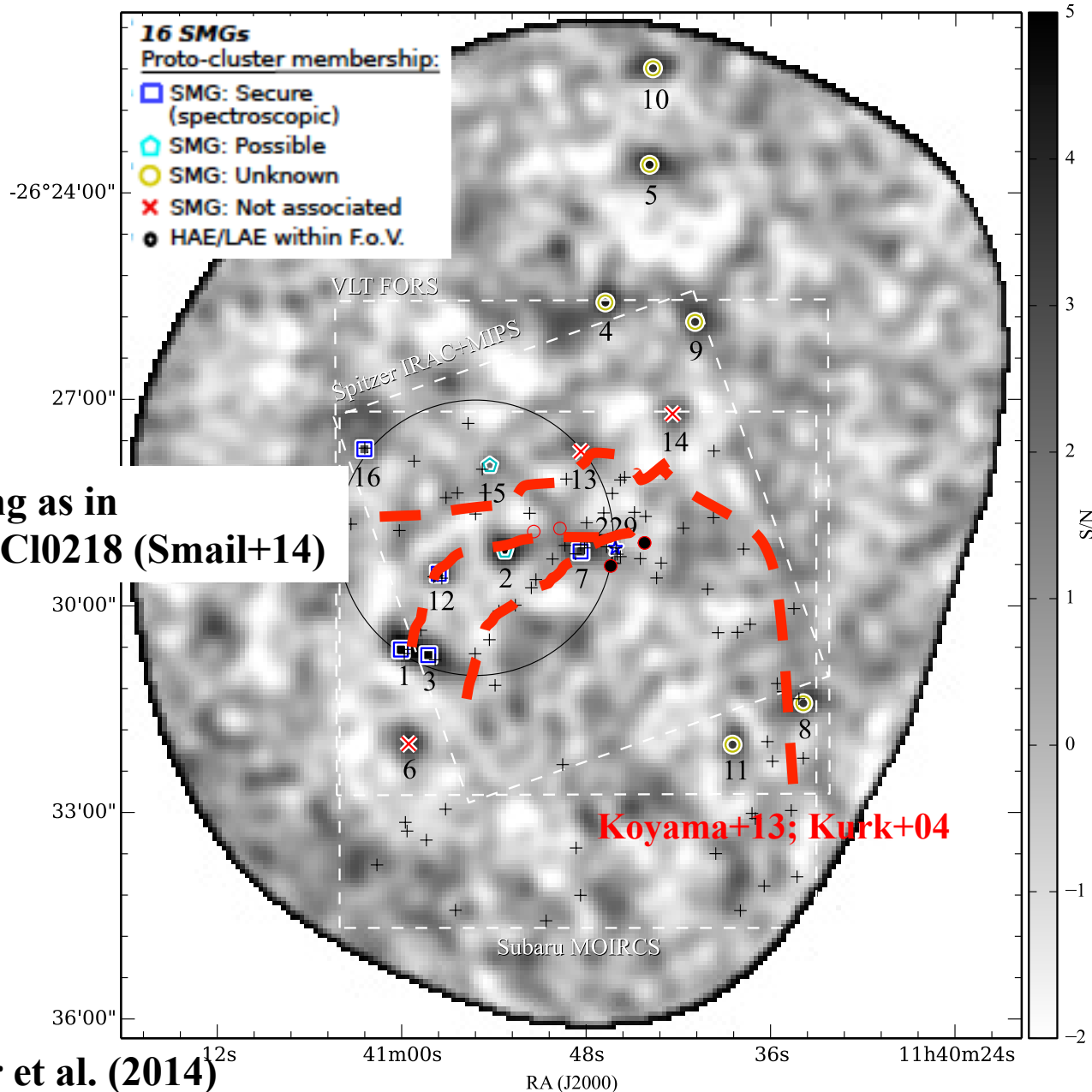


# Proto-Cluster Membership

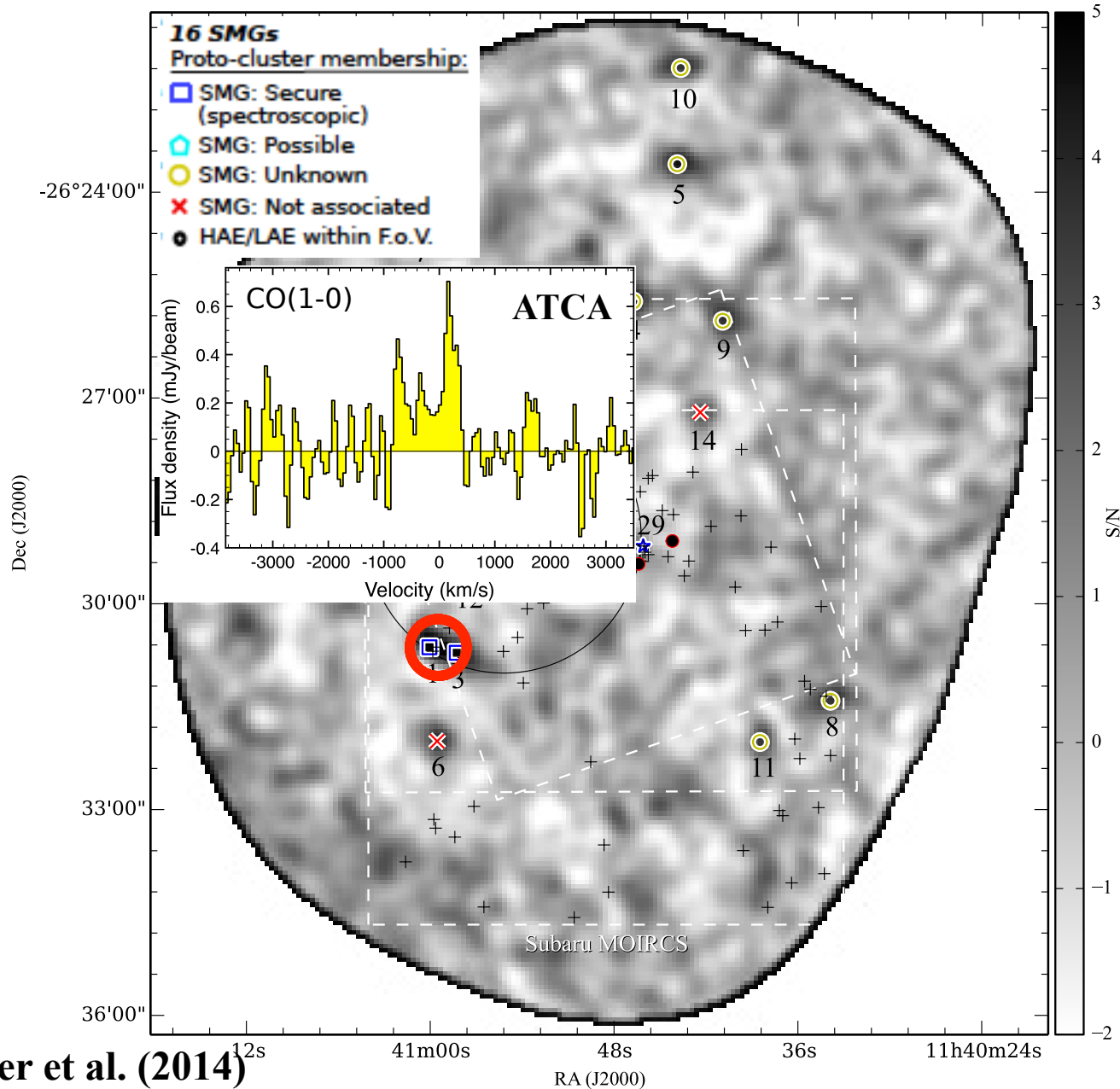




# Proto-Cluster Membership



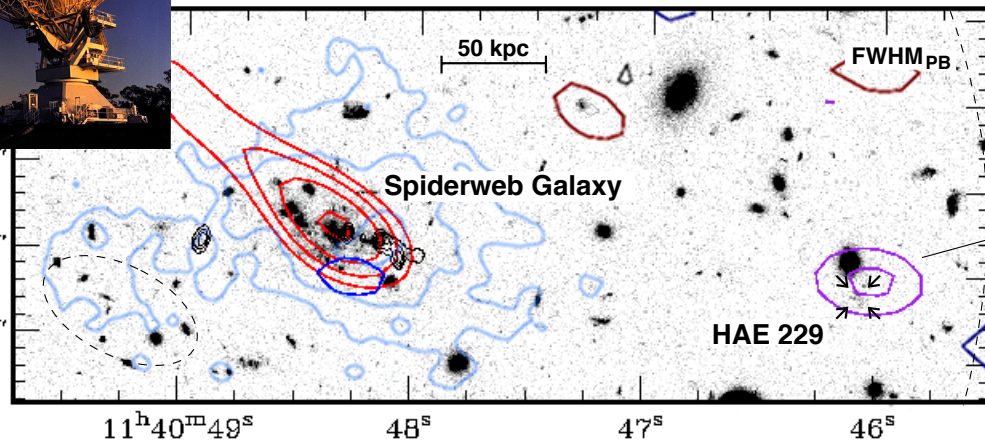
# Proto-Cluster Membership



# HAE229 .... alias SCUBA SMG#3



Declination (

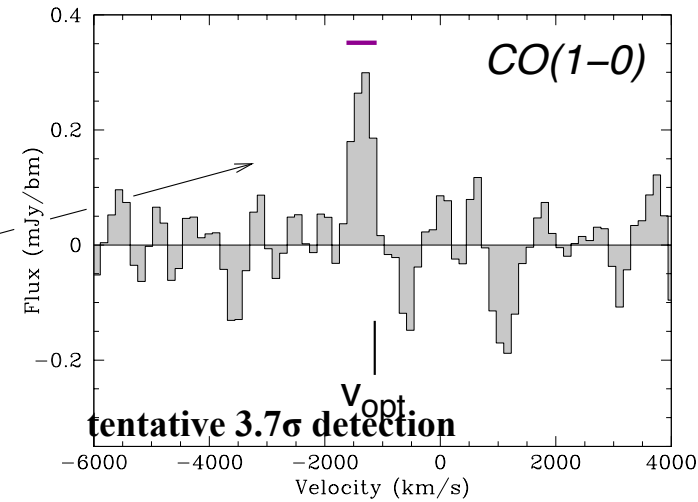
 -29°05"  
 -29°10"  
 -29°15"

 11<sup>h</sup>40<sup>m</sup>49<sup>s</sup>

 48<sup>s</sup>

 47<sup>s</sup>

 46<sup>s</sup>

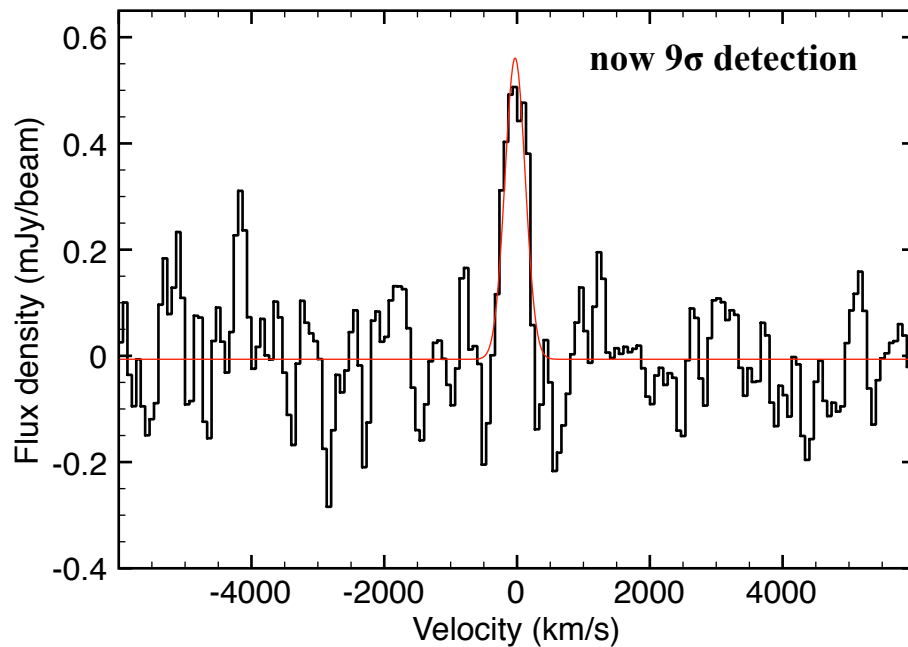
Right Ascension (J2000)



tentative 3.7σ detection

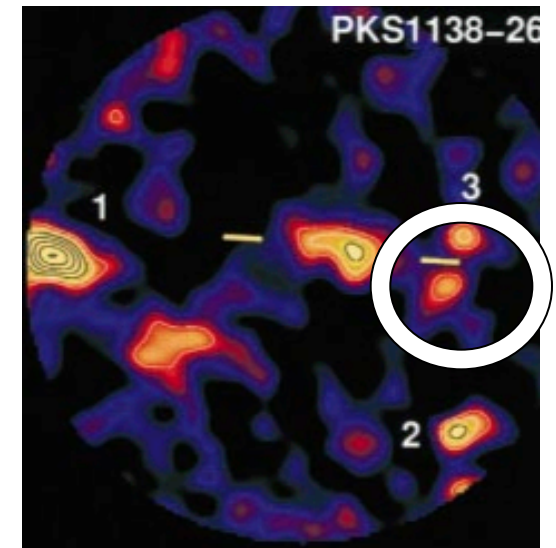
Velocity (km/s)

Emonts et al. (2013)



now 9σ detection

Velocity (km/s)



PKS1138-26

1

2

3

Dannerbauer, Emonts et al. to be subm.

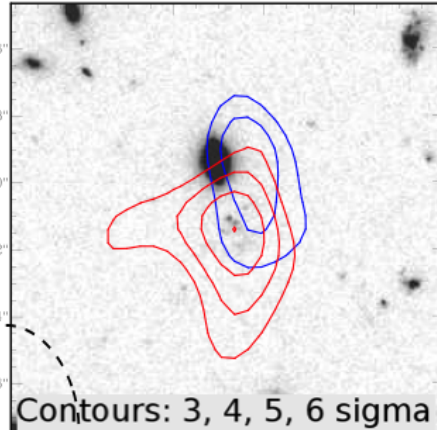
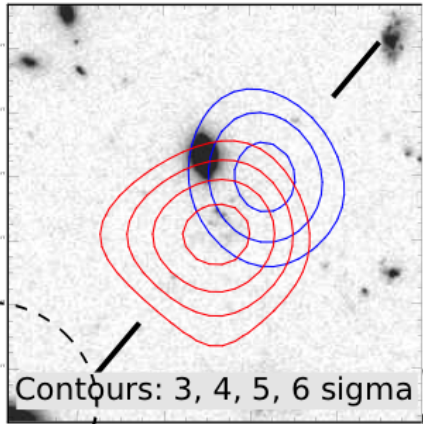


# HAE229

## Main sequence galaxy

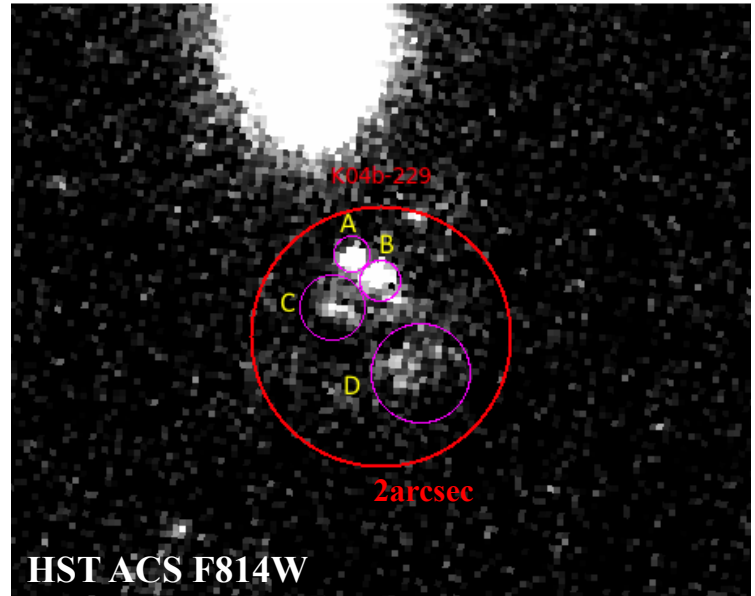
uvrange = 0 - 31 k-lambda

uvrange = 0 - 60 k-lambda



Contours: 3, 4, 5, 6 sigma

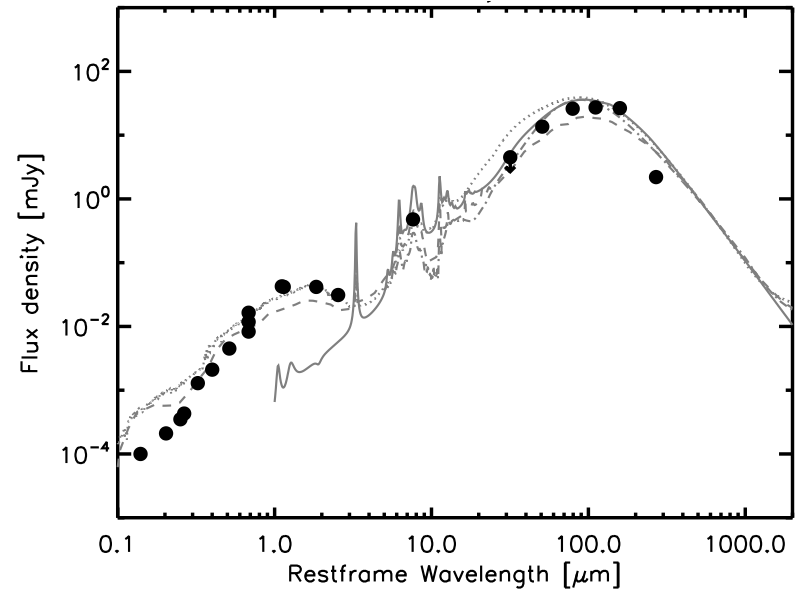
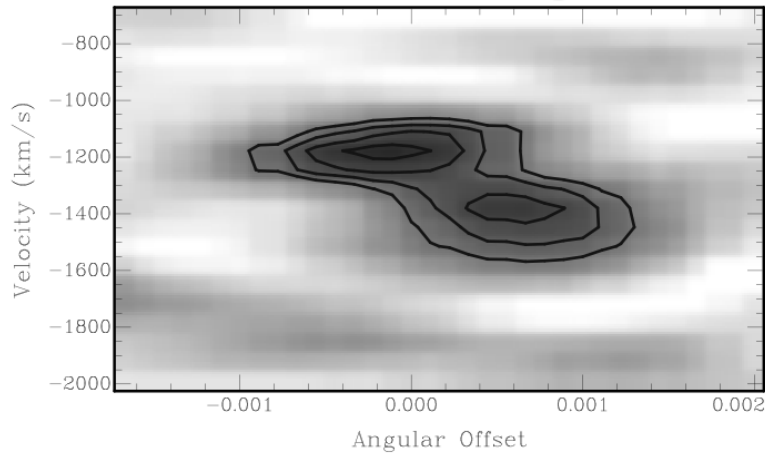
Contours: 3, 4, 5, 6 sigma



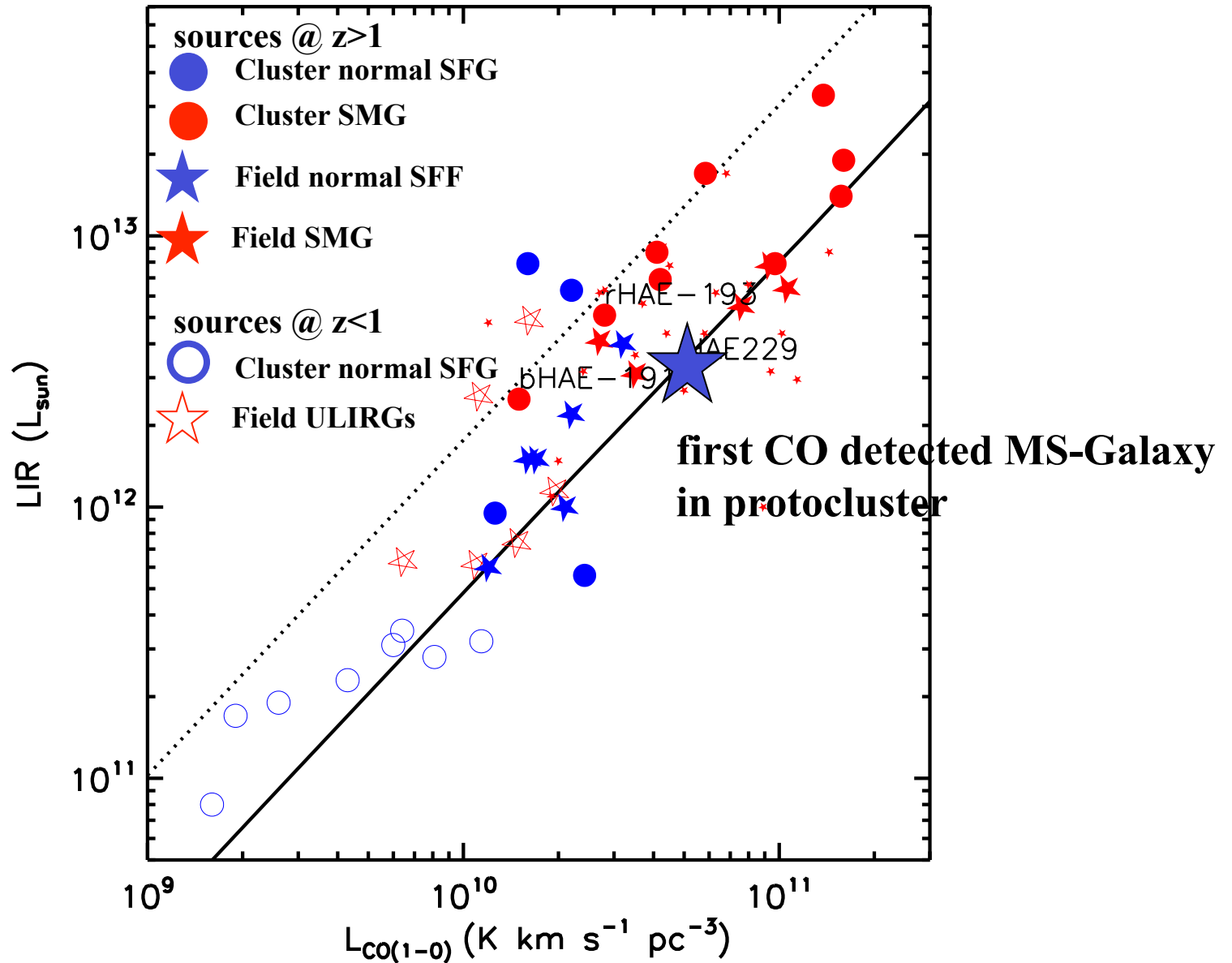
HST ACS F814W

HAE229

Contours: 3, 4, 5, 6 sigma



# HAE229



# Conclusion

- **overdensity of dusty starbursts associated to MRC1138**
- **at least 8 SMGs are members of the proto-cluster**
- **HAEs promising approach to trace (part of) the SMG (system)**
- **LABOCA (submm) imaging seems to be a good way to search for overdensities/large scale structures → Poster #28 Umehata see also for  $z > 1$  clusters: #27 Noble, #46 Webb and #48 Santos**
- **first CO of main-sequence galaxy in protocluster**
  - *Dannerbauer, Kurk, De Breuck et al., 2014, A&A, 2014, 570, 55*
  - *Dannerbauer, Emonts et al., 2015, to be submitted*